



Quantum Physics Interface to Data Science, Artificial Intelligence & Spacekime Analytics

Ivo D. Dinov


Statistics Online Computational Resource
Health Behavior & Biological Sciences
Computational Medicine & Bioinformatics
Michigan Institute for Data Science
University of Michigan

<https://SOCR.umich.edu>


Joint work with Milen V. Velev (BTU) & Yueyang Shen (UM)

Based on the book "Data Science: Time Complexity, Inferential Uncertainty & Spacekime Analytics"



STATISTICS ONLINE COMPUTATIONAL RESOURCE (SOCR)
UNIVERSITY OF MICHIGAN



Slides Online:
"SOCR News"



Outline


- ❑ Complex-Time (*kime*) & Rationale
- ❑ Solutions of untrahyperbolic wave equations
- ❑ Open Spacekime Problems

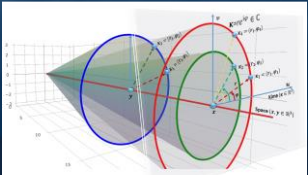
- ❑ Data/Neuro Science Applications
 - ❑ Random Sampling vs. Hidden Variables Paradigm
 - ❑ Neuroimaging (fMRI): time-series → kime-surfaces
- ❑ Bayesian Formulation of Spacekime Inference
- ❑ Live Demo Links

Complex-Time (*Kime*)

- ❑ At a given spatial location, \mathbf{x} , complex time (*kime*) is defined by $\kappa = r e^{i\varphi} \in \mathbb{C}$, where:
 - ❑ the magnitude represents the longitudinal events order ($r > 0$) and characterizes the longitudinal displacement in time, and
 - ❑ event phase ($-\pi \leq \varphi < \pi$) is an angular displacement, or event direction
- ❑ There are multiple alternative parametrizations of kime in the complex plane
- ❑ Space-kime manifold is $\mathbb{R}^3 \times \mathbb{C}$:
 - ❑ (x, k_1) and (x, k_4) have the same spacetime representation, but different spacekime coordinates,
 - ❑ (x, k_1) and (y, k_1) share the same kime, but represent different spatial locations,
 - ❑ (x, k_2) and (x, k_3) have the same spatial-locations and kime-directions, but appear sequentially in order, $t_2 < t_1$.






Rationale for *Time* → *Kime* Extension

- ❑ **Math** – *Time* is a special case of *kime*, $\kappa = |\kappa| e^{i\varphi}$ where $\varphi = 0$ (nil-phase)
 - algebraically a *multiplicative* (algebraic) group, (multiplicative) unity (identity) = 1
 - multiplicative inverses, multiplicative identity, associativity $t_1 * (t_2 * t_3) = (t_1 * t_2) * t_3$
 - The *time* domain (\mathbb{R}^+) is not a complete algebraic field $(+, *)$:
 - Additive unity (0), element additive inverse $(-t)$: $t + (-t) = 0$; is outside \mathbb{R}^+ (time-domain)
 - $x^2 + 1 = 0$ has no solutions in time (or in \mathbb{R}) ...

$$\text{Group}(*) \subseteq \text{Ring} \left(\begin{array}{c} \text{Compatible operations} \\ (+, *) \\ \text{associative \& distributive} \end{array} \right) \subseteq \text{Field} \left(\begin{array}{c} \text{Group}(+) \\ (+, *) \end{array} \right)$$

- Classical time (\mathbb{R}^+) is a *positive cone* over the field of the real numbers (\mathbb{R})
- Time forms a subgroup of the multiplicative group of the reals
- Whereas kime (\mathbb{C}) is an algebraically *closed prime field* that naturally extends time
- *Time* is ordered & *kime* is not – the kime magnitude preserves the intrinsic time order
- Kime (\mathbb{C}) represents the smallest natural extension of time, complete field that agrees with time
- The *time* group is closed under addition, multiplication, and division (but not subtraction). It has the topology of \mathbb{R} and the structure of a multiplicative topological group \cong additive topological semigroup

- ❑ **Physics** –
 - ❑ Problem of time ... (DOI: 10.1007/978-3-319-58848-3)
 - ❑ \mathbb{R} and \mathbb{C} Hilbert-space quantum theories make different predictions (DOI: 10.1038/s41586-021-04160-4)
- ❑ **AI/Data Science** – Random IID sampling, Bayesian reps, tensor modeling of \mathbb{C} kimesurfaces, novel analytics



Ultrahyperbolic Wave Equation – Cauchy Initial Data

❑ **Nonlocal constraints** yield the existence, uniqueness & stability of local and global solutions to the ultrahyperbolic wave equation under Cauchy initial data ...

$$\sum_{i=1}^{d_s} \partial_{x_i}^2 u \equiv \Delta_x u(x, \kappa) = \Delta_\kappa u(x, \kappa) \equiv \sum_{i=1}^{d_t} \partial_{\kappa_i}^2 u, \quad \begin{cases} u|_0 = u(x, 0, \kappa_{-1}) = f(x, \kappa_{-1}) \\ u_1 = \partial_{\kappa_1} u(x, 0, \kappa_{-1}) = g(x, \kappa_{-1}) \end{cases}$$


spatial Laplacian
temporal Laplacian
initial conditions (Cauchy Data)

where $\mathbf{x} = (x_1, x_2, \dots, x_{d_s}) \in \mathbb{R}^{d_s}$ and $\kappa = (\kappa_1, \kappa_2, \dots, \kappa_{d_t}) \in \mathbb{R}^{d_t}$ are the Cartesian coordinates in the d_s space and d_t time dims. Stable local solution over a Fourier frequency region defined by **nonlocal constraints** $|\xi| \geq |\eta_{-1}|$:

$$\hat{u}(\xi, \kappa_1, \eta_{-1}) = \cos(2\pi \kappa_1 \sqrt{|\xi|^2 - |\eta_{-1}|^2}) \hat{u}_0(\xi, \eta_{-1}) + \sin(2\pi \kappa_1 \sqrt{|\xi|^2 - |\eta_{-1}|^2}) \frac{\hat{u}_1(\xi, \eta_{-1})}{2\pi \sqrt{|\xi|^2 - |\eta_{-1}|^2}}$$

where $\mathcal{F} \left(\begin{smallmatrix} u_0 \\ u_1 \end{smallmatrix} \right) = \left(\begin{smallmatrix} \hat{u}_0 \\ \hat{u}_1 \end{smallmatrix} \right) = \left(\begin{smallmatrix} \hat{u}_0(\xi, \eta_{-1}) \\ \hat{u}_1(\xi, \eta_{-1}) \end{smallmatrix} \right) = \left(\begin{smallmatrix} \hat{u}(\xi, \eta_{-1}) \\ \partial_{\kappa_1} \hat{u}(\xi, \eta_{-1}) \end{smallmatrix} \right)$

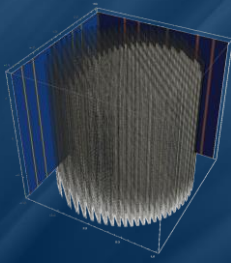

$$u(x, \kappa_1, \kappa_{-1}) = \mathcal{F}^{-1}(\hat{u})(x, \kappa) = \int_{D_s \times D_{t-1}} \hat{u}(\xi, \kappa_1, \eta_{-1}) \times e^{2\pi i(x\xi)} \times e^{2\pi i(\kappa_1 \eta_{-1})} d\xi d\eta_{-1}$$



Wang et al., 2022 | Dinov & Velev (2021)

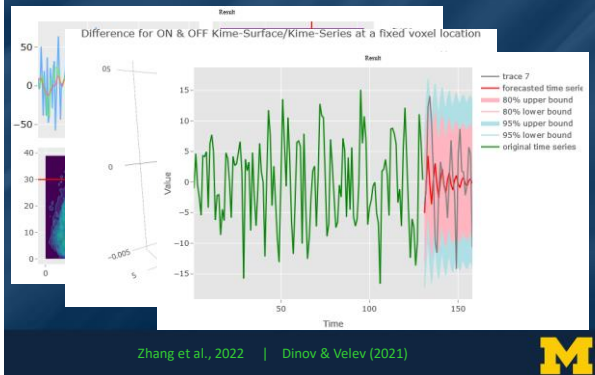
A Spacekime Solution to Wave Equation

- ❑ **Math Generalizations:**
 - Derived other spacekime concepts: law of addition of velocities, energy-momentum conservation law, stability conditions for particles moving in spacekime, conditions for nonzero rest particle mass, causal structure of spacekime, and solutions of the ultrahyperbolic wave equation under Cauchy initial data ...

Wang et al., 2022 | Dinov & Velev (2021)

Spacetime Time-series \Rightarrow Spacekime Kimesurfaces \Rightarrow TLM



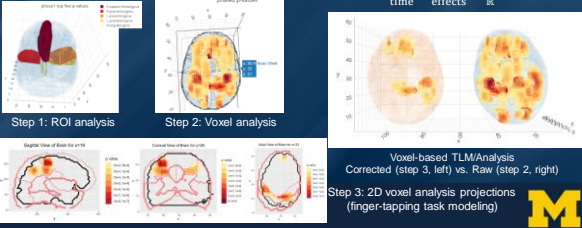
Zhang et al., 2022 | Dinov & Velev (2021)



Tensor-based Linear Modeling of fMRI

3-Step Analysis: registering the fMRI data into a brain atlas space, 56 ROIs, tensor linear modeling, post-hoc FDR processing & selection of large clusters of significant voxels are identified within the important ROIs: $Y = \overbrace{(X, B)}^{\text{tensor product}} + E$

The dimensions of the tensor Y are $160 \times a \times b \times c$, where the tensor elements represent the response variable $Y[t, x, y, z]$, i.e., fMRI intensity. For fMRI magnitude (real-valued signal), the design tensor X dimensions are: $160 \times \frac{4}{\text{time}} \times \frac{1}{\text{effects}} \times \frac{1}{\text{R}}$



Bayesian Inference Representation

- Suppose we have a single spacetime observation $X = \{x_{i_0}\} \sim p(x | \gamma)$ and $\gamma \sim p(\gamma | \varphi = \text{phase})$ is a process parameter (or vector) that we are trying to estimate.
- Spacekime analytics aims to make appropriate inference about the process X .
- The sampling distribution, $p(x | \gamma)$, is the distribution of the observed data X conditional on the parameter γ and the prior distribution, $p(\gamma | \varphi)$, of the parameter γ before the data X is observed, $\varphi = \text{phase aggregator}$.
- Assume that the hyperparameter (vector) φ , which represents the kime-phase estimates for the process, can be estimated by $\hat{\varphi} = \varphi'$.
- Such estimates may be obtained from an oracle (model distribution), approximated using similar datasets, acquired as phases from samples of analogous processes, derived via some phase-aggregation strategy, or computed via Laplace transform.
- Let the posterior distribution of the parameter γ given the observed data $X = \{x_{i_0}\}$ be $p(\gamma | X, \varphi')$ and the process parameter distribution of the kime-phase hyperparameter vector φ be $\gamma \sim p(\gamma | \varphi)$.



Bayesian Inference Simulation

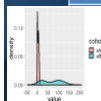
- Simulation example using 2 random samples drawn from mixture distributions each of $n_A = n_B = 10K$ observations:
 - $\{X_{A,i}\}_{i=1}^{10K}$, where $X_{A,i} = 0.3U_i + 0.7V_i$, $U_i \sim N(0,1)$ and $V_i \sim N(5,3)$, and
 - $\{X_{B,i}\}_{i=1}^{10K}$, where $X_{B,i} = 0.4P_i + 0.6Q_i$, $P_i \sim N(20,20)$ and $Q_i \sim N(100,30)$.
- The intensities of cohorts A and B are independent and follow different mixture distributions. We'll split the first cohort (A) into training (C) and testing (D) subgroups, and then:
 - Transform all four cohorts into Fourier k-space,
 - Iteratively randomly sample single observations from the (training) cohort C ,
 - Reconstruct the data into spacetime using a single kime-magnitude value and alternative kime-phase estimates derived from cohorts B , C , and D , and
 - Compute the classical spacetime-derived population characteristics of cohort A and compare them to their spacekime counterparts obtained using a single C kime-magnitude paired with B , C , or D kime-phases.



Bayesian Inference Simulation

Summary statistics for the original process (cohort A) and the corresponding values of their counterparts computed using the spacekime reconstructed signals based on kime-phases of cohorts B , C , and D . The estimates for the latter three cohorts correspond to reconstructions using a single spacetime observation (i.e., single kime-magnitude) and alternative kime-phases (in this case, kime-phases derived from cohorts B , C , and D).

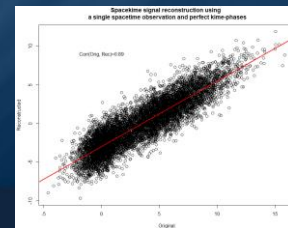
	Spacetime (A)	Spacekime Reconstructions (single kime-magnitude)		
Summaries	Original	(B) Phase=Diff. Process	(C) Phase=True	(D) Phase-independent
Min	-2.38798	-3.798440	-2.98116	-2.69808
1 st Quartile	-0.89359	-0.636799	-0.76765	-0.76453
Median	0.03311	0.009279	-0.05982	-0.08329
Mean	0.00000	0.000000	0.00000	0.00000
3 rd Quartile	0.75772	0.645119	0.72795	0.69889
Max	3.61346	3.986702	3.64800	3.22987
Skewness	0.348269	0.001021943	0.2372526	0.31398
Kurtosis	-0.68176	0.2149918	-0.4452207	-0.3270084



Bayesian Inference Simulation

The correlation between the original data (A) and its reconstruction using a single kime magnitude and the correct kime-phases (C) is $\rho(A, C) = 0.89$.

This strong correlation suggests that a substantial part of the A process energy can be recovered using only a single observation. In this case, to reconstruct the signal back into spacetime and compute the corresponding correlation, we used a single kime-magnitude (sample-size=1) and process C kime-phases.



Bayesian Inference Simulation

Let's demonstrate the Bayesian inference corresponding to this spacetime data analytic problem using a simulated bimodal experiment:

$$X_A = 0.3U + 0.7V, \text{ where } U \sim N(0,1) \text{ and } V \sim N(5,3)$$

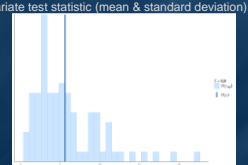
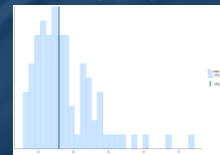
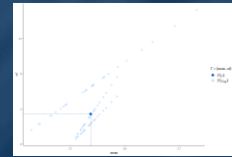
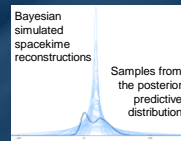
Specifically, we will illustrate the Bayesian inference using repeated single spacetime observations from cohort A , $X = \{x_{i,j}\}$, and varying kime-phase priors ($\theta =$ phase aggregator) obtained from cohorts B , C , or D , using different posterior predictive distributions.

Relations between the empirical data distribution (dark blue) and samples from the posterior predictive distribution, representing Bayesian simulated spacetime reconstructions (light-blue). The derived Bayesian estimates do not perfectly match the empirical distribution of the simulated data, yet there is clearly information encoding that is captured by the spacetime data reconstructions.

This signal compression can be exploited by subsequent model-based or model-free data analytic strategies for retrospective prediction, prospective forecasting, ML classification, AI derived clustering, and other spacetime inference methods.



Bayesian Inference Simulation



Test statistic (maximum)

Test statistic (inter-quartile range, IQR)

Relations between the empirical data distribution (dark blue) and samples from the posterior predictive distribution, Bayesian simulated spacetime reconstructions (light-blue).



Spacekime Analytics: Demos

Tutorials

- <https://TCIU.predictive.space>
- <https://SpaceKime.org>

R Package

- <https://cran.rstudio.com/web/packages/TCIU>

GitHub

- <https://github.com/SOCR/TCIU>



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Collaborators

- SOCR: Milen Velev, Yueyang Shen, Daxuan Deng, Zijing Li, Yongkai Qiu, Zhe Yin, Yufei Yang, Yuxin Wang, Rongqian Zhang, Yuyao Liu, Yupeng Zhang, Yunjie Guo, Simeone Marino
- UMich MIDAS/MCAIM Centers: Lydia Bieri, Kayvan Najarian, Chris Monk, Issam El Naqa, HV Jagadish, Brian Athey



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